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FROM JET QUENCHING TO WAVE TURBULENCE

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We discuss average properties of the gluon cascade generated by an energetic parton propagating through a dense QCD medium. The cascade is mostly made with relatively soft gluons, whose production is not suppressed by the LPM effect. Unlike for usual QCD cascades in the vacuum, where the typical splittings are very asymmetric (soft and collinear), the medium-induced branchings are quasi-democratic and lead to wave turbulence. This results in a very efficient mechanism for the transport of energy at large angles with respect to the jet axis, which might explain the di-jet asymmetry observed in Pb–Pb collisions at the LHC.

One remarkable phenomenon discovered in the lead–lead collisions at the LHC is the *di-jet asymmetry*, a strong imbalance between the energies of two back–to–back jets. This asymmetry is commonly attributed to the effect of the interactions of one of the two jets with the dense QCD matter that it traverses, while the other leaves the system unaffected. Originally identified [1,2] as missing energy, this phenomenon has been subsequently shown [3] to consist in the transport of a sizable part (about 10%) of the total jet energy by *soft particles* towards *large angles*. The total amount of energy which is thus lost by the jet, ~ 20 GeV, is considerably larger than the typical transverse momentum, ~ 1 GeV, of a parton in the medium. In that sense, the effect is large and potentially non-perturbative. Yet, there exists a mechanism within perturbative QCD which can naturally explain the energy loss at large angles: the BDMPSZ mechanism for medium-induced gluon radiation (from Baier, Dokshitzer, Mueller, Peigné, Schiff [4] and Zakharov [5]). Most previous studies within this approach have focused on the energy lost by the leading particle, which is controlled by rare and relatively hard emissions at small angles. More recently, in the wake of the LHC data, the attention has been shifted towards softer emissions, for which the effects of *multiple branching* become important. The generalization of the BDMPSZ formalism to multiple branchings has only recently been given [6,7], and this turns out to have interesting physical consequences [8]. These recent developments will be briefly reviewed in what follows.

The BDMPSZ mechanism relates the radiative energy loss by an energetic parton propagating through a dense QCD medium (‘quark–gluon plasma’) to the transverse momentum broadening via scattering off the medium constituents. A central concept is the *formation time* $\tau_f(\omega)$ — the typical times it takes a gluon with energy $\omega \ll E$ to be emitted. (E is the energy of the original parton, a.k.a. the ‘leading particle’.) The gluon starts as a virtual fluctuation which moves away from its parent parton via quantum diffusion: the transverse^a separation b_\perp grows with time as $b_\perp^2 \sim \Delta t/\omega$. The gluon can be considered as ‘formed’ when it loses coherence w.r.t to its source, meaning that b_\perp is at least as large as the gluon transverse wavelength

^aThe ‘transverse directions’ refer to the 2-dimensional plane orthogonal to the 3-momentum of the leading particle, which defines the ‘longitudinal axis’.

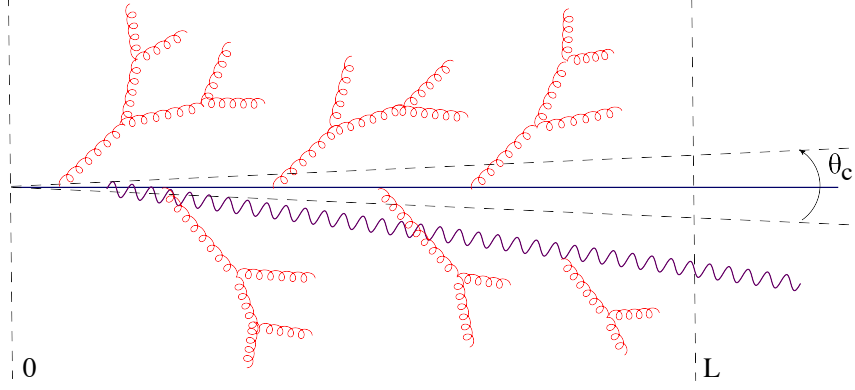


Figure 1: The jet produced in a ‘rare event’ which, besides the leading particle and the (quasi-deterministic) soft gluon cascades at large angles, also contains a hard gluon with $\omega \sim \omega_c$, which propagates at a small angle $\theta \sim \theta_c$.

$\lambda_\perp = 1/k_\perp$. But the gluon transverse momentum k_\perp is itself increasing with time, via collisions which add random kicks Δk_\perp at a rate given by the *jet quenching parameter* \hat{q} : $\Delta k_\perp^2 \sim \hat{q} \Delta t$. The ‘formation’ condition, $b_\perp \gtrsim 1/\Delta k_\perp$ for $\Delta t \gtrsim \tau_f$, implies (below, k_f and θ_f are the typical values of the gluon transverse momentum and emission angle at the time of formation)

$$\tau_f(\omega) \simeq \sqrt{\frac{2\omega}{\hat{q}}}, \quad k_f^2 = \hat{q} \tau_f(\omega) \simeq (2\omega \hat{q})^{1/2}, \quad \theta_f \simeq \frac{k_f}{\omega} \simeq \left(\frac{2\hat{q}}{\omega^3} \right)^{1/4}. \quad (1)$$

Eq. (1) applies as long as $\ell \ll \tau_f(\omega) < L$, where L is the length of the medium and ℓ is the mean free path between successive collisions. The second inequality implies an upper limit on the energy of a gluon that can be emitted via this mechanism, and hence a lower limit on the emission angle: $\omega \lesssim \omega_c \equiv \hat{q} L^2/2$ and $\theta_f \gtrsim \theta_c \equiv 2/(\hat{q} L^3)^{1/2}$. The BDMPSZ regime corresponds to $\hat{q} L^3 \gg 1$ and hence $\theta_c \ll 1$. Choosing $\hat{q} = 1 \text{ GeV}^2/\text{fm}$ (the weak coupling estimate [4]) for a QGP with temperature $T = 250 \text{ MeV}$ and $L = 4 \text{ fm}$, one finds $\omega_c \simeq 40 \text{ GeV}$ and $\theta_c \simeq 0.05$.

Eq. (1) shows that the relatively soft gluons with $\omega \ll \omega_c$ have (i) short formation times $\tau_f(\omega) \ll L$ and (ii) large formation angles $\theta_f \gg \theta_c$. Both properties are important for us here. Property (i) implies that such gluons are produced *abundantly*: their emission can be initiated at any place inside the medium, hence the associated spectrum (below, $\bar{\alpha} \equiv \alpha_s N_c/\pi$)

$$\omega \frac{dN}{d\omega} \simeq \frac{\alpha_s N_c}{\pi} \frac{L}{\tau_f(\omega)} = \bar{\alpha} \sqrt{\frac{\omega_c}{\omega}}, \quad (2)$$

is enhanced by a factor $L/\tau_f(\omega) \gg 1$, which expresses the relative longitudinal phase-space available for their emission. Property (ii) shows that the soft gluons have the potential to transport a part of the jet energy towards large angles. This looks like a small effect, since soft gluons carry only little energy, but this is enhanced by multiple emissions, as we shall see.

Eq. (2) is the BDMPSZ spectrum for a soft, medium-induced, gluon emission [4,5]. Note that the emission probability is small, of $\mathcal{O}(\bar{\alpha})$, for a relatively hard gluon with $\omega \sim \omega_c$. This is a consequence of the *LPM effect* (from Landau, Pomeranchuk, Migdal) — the fact that one needs a large number $\tau_f(\omega)/\ell \gg 1$ of successive scatterings in order to produce a single gluon. Such a rare but hard emission dominates the average energy loss by the leading particle, $\Delta E \equiv \int^{\omega_c} d\omega \omega (dN/d\omega) \sim \bar{\alpha} \omega_c$, but it cannot contribute to the observed di-jet asymmetry, because a hard gluon propagates at a very small angle $\sim \theta_c$ w.r.t. the jet axis.

To understand di-jet asymmetry, one rather needs to focus on soft gluon emissions at large angles. For such emissions, the effects of *multiple branching* become important, as we now explain. The probability for a single emission, as measured by the spectrum (2), exceeds unity

^bOne can estimate $\hat{q} \simeq m_D^2/\ell$, where m_D is the Debye screening mass in the medium. Eq. (1) applies to a medium whose size L is much larger than both the screening length $1/m_D$ and the mean free path ℓ .

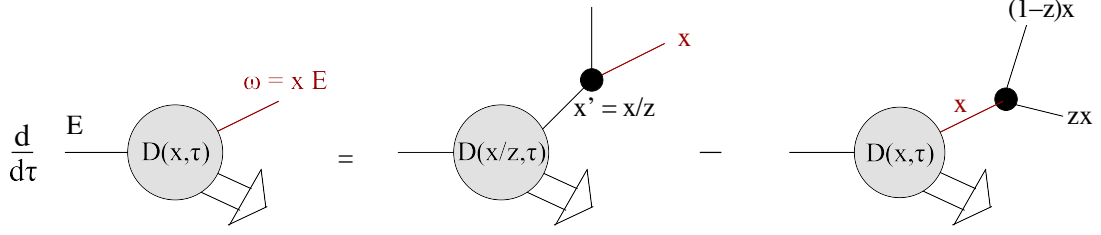


Figure 2: The change in the gluon spectrum $D(x, \tau) \equiv x(dN/dx)$ due to one additional branching $g \rightarrow gg$.

when $\omega \lesssim \bar{\alpha}^2 \omega_c$. In this non-perturbative regime at small ω , Eq. (2) must be corrected to account for multiple emissions by the leading particle, and also for the subsequent branching of the soft primary gluons into even softer ones (thus leading to gluon cascades; see Fig. 1). The soft branchings are quasi-deterministic and can be observed in an event-by-event basis.

The treatment of multiple branching is *a priori* complicated by interference effects between emissions from different partonic sources. For the case of a jet evolving into the vacuum, such effects are known to lead to *angular ordering* between the successive emissions. Interference effects for medium-induced gluon radiation started to be investigated only recently [6,9,10,7]. The respective analysis is quite involved, but its main conclusion is very simple [7]: the interference effects for soft emissions of the BDMPSZ type are *negligible*, since suppressed by a factor $\tau_f(\omega)/L \ll 1$. To understand this result, recall that, in order to interfere with each other, two emitters must be *coherent* with each other, a situation which can occur if they have a common ancestor. However, the partons produced by a medium-induced branching lose their mutual *colour coherence* already during the formation process, because they randomly scatter in the medium. Accordingly, they can interfere with each other only during a short time $\tau_f(\omega)$, which gives a small phase-space whenever $\omega \ll \omega_c$. This implies that successive medium-induced emissions can be considered as *independent* of each other and taken into account via a *probabilistic branching process*, in which the BDMPSZ spectrum plays the role of the elementary branching rate^c.

The general branching process is a Markovian process in 3+1 dimensions which describes the gluon distribution in energy (ω) and transverse momentum (k_\perp), and its evolution when increasing the medium size L . This process is well suited for numerical studies via Monte-Carlo simulations. But analytic results have also been obtained [8], for a simplified process in 1+1 dimensions, which describes the energy distribution alone. These results lead to an interesting physical picture that we shall now describe.

To that aim, it is convenient to focus on the *gluon spectrum* $D(x, \tau) \equiv x(dN/dx)$, where $x \equiv \omega/E$ is the energy fraction carried by a gluon from the jet and the ‘evolution time’ τ is the medium size in dimensionless units, as defined in Eq. (3) below. The quantity $D(x, \tau)$ obeys an evolution equation [8,11,13], which is formally similar to the DGLAP equation describing the fragmentation of a jet in the vacuum: a rate equation, which involves a ‘gain’ term and a ‘loss’ term and is illustrated in Fig. 2. The ‘gain’ term describes the increase in the number of gluons with a given x via radiation from gluons with a larger $x' = x/z$, with any $x < z < 1$. The ‘loss’ term expresses the decrease in the number of gluons at x via their decay $x \rightarrow zx, (1-z)x$, with any $0 < z < 1$. What is specific to the problem at hand, is the particular form of the in-medium branching rate — the BDMPSZ spectrum in Eq. (2) —, which is very different from the DGLAP splitting function. This difference has important physical consequences, that can be best appreciated by comparing the respective solution $D(x, \tau)$ to the BDMPSZ spectrum for a single emission and to the DGLAP spectrum for gluon evolution in the vacuum.

By construction, the BDMPSZ spectrum coincides with the first iteration to the evolution

^cSuch a classical branching process, obtained by iterating the single BDMPSZ emission, has already been used in applications to phenomenology, albeit on a heuristic basis [11,12,13].

equation. For soft gluons ($x \ll 1$), it is given by Eq. (2), which in our new notations reads

$$D^{(1)}(x \ll 1, \tau) \simeq \frac{\tau}{\sqrt{x}}, \quad \text{with} \quad \tau \equiv \bar{\alpha} \sqrt{\frac{\hat{q}}{E}} L. \quad (3)$$

This spectrum increases quite fast with $1/x$ and thus predicts that a non-negligible fraction of the total radiated energy is emitted directly at large angles: the energy fraction transported via a single gluon emission at angles larger than a given value θ_0 is estimated as

$$\mathcal{E}^{(1)}(\theta > \theta_0, \tau) = \int_0^{x_0} dx D^{(1)}(x, \tau) \simeq 2\tau\sqrt{x_0} \quad \text{with} \quad \theta_0 \simeq \left(\frac{2\hat{q}}{x_0^3 E^3} \right)^{1/4}. \quad (4)$$

This quantity rapidly decreases with increasing θ_0 (i.e. with decreasing x_0), showing that direct radiation by the leading particle cannot explain the large energy loss at large angles observed in relation with the di-jet asymmetry at the LHC [3]. As we shall shortly argue, multiple branching provides a much more efficient mechanism in that sense.

The approximation in Eqs. (3)–(4) breaks down when $D^{(1)}(x, \tau) \sim \mathcal{O}(1)$, meaning for $x \lesssim \tau^2$. In this non-perturbative regime at small x , one needs an exact result which resums multiple branching to all orders. Before we present this result, it is instructive to summarize the picture one would expect on the basis of our experience with other parton cascades in QCD, like DGLAP. (These expectations turn out to be naive, but their failure will be instructive.)

Via successive branchings, the partons at large x get replaced via partons with smaller values of x , which must be numerous enough to carry the energy of their parents. This seems to imply that the rise in the gluon distribution $D(x, \tau)$ at small x must become steeper and steeper with increasing τ , in such a way to accommodate the energy which disappears at larger values of x . However, this last conclusion is based on the tacit assumption that ‘the energy remains in the spectrum’, meaning that the energy sum-rule $\int_0^1 dx D(x, \tau) = 1$ is satisfied at any τ . If that was the case, this would also impose a strong limitation on the energy that can be carried by the small- x gluons for a given value of τ : in order for the function $D(x, \tau)$ to be integrable as $x \rightarrow 0$, the integral $\mathcal{E}(x_0, \tau) \equiv \int_0^{x_0} dx D(x, \tau)$ must vanish as a power of x_0 when $x_0 \rightarrow 0$, meaning that the energy fraction radiated at large angles could not be too large. Such a scenario would have little chance to explain the LHC data for di-jet asymmetry. However, this is *not* the picture that emerges from the in-medium evolution equation and that we now describe.

As mentioned, an exact analytic solution is known, modulo some harmless simplifications in the branching rate [8]. This is shown here only for $x \ll 1$, where it reads

$$D(x \ll 1, \tau) \simeq \frac{\tau}{\sqrt{x}} e^{-\pi\tau^2}. \quad (5)$$

(The global spectrum for any x is illustrated in Fig. 3 for various values of τ .) The spectrum (5) is remarkable in several aspects: (i) The ‘scaling’ behaviour at small x , $D(x) \propto 1/\sqrt{x}$, is the same as for the BDMPSZ spectrum, Eq. (3). Formally, one can read Eq. (5) as ‘direct radiation by the leading particle \times survival probability for the latter’. However, unlike Eq. (3), the spectrum (5) includes the effects of multiple branching to all orders. This demonstrates that the scaling spectrum is a *fixed point* of the evolution, for which the ‘gain’ term and the ‘loss’ term precisely cancel each other. (ii) The fact that the shape of the spectrum at $x \ll 1$ does not change with τ means that its small- x region cannot accommodate the energy which disappears via splittings at large x . This is also visible in Fig. 3: at small $\tau \ll 1/\sqrt{\pi}$, the small- x spectrum increases linearly with τ , as the BDMPSZ spectrum (3). At the same time, the leading-particle peak, which originally was at $x = 1$, moves towards $x < 1$ and gets broader. For larger times $\tau \gtrsim 1/\sqrt{\pi}$, the spectrum is globally suppressed by the Gaussian factor in (5). Clearly, the energy disappears from the spectrum with increasing τ , as confirmed by an explicit calculation of the energy sum-rule [8]: $\int_0^1 dx D(x, \tau) = e^{-\pi\tau^2}$. Where does the energy go?

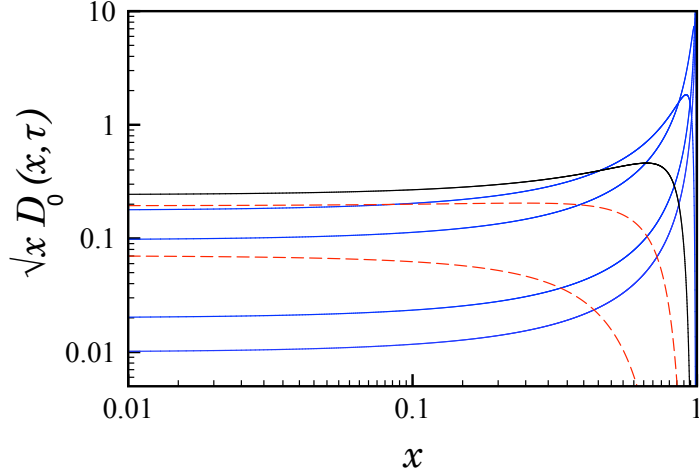


Figure 3: Plot (in Log-Log scale) of $\sqrt{x}D(x, \tau)$ as a function of x for various values of τ (full lines from bottom to top: $\tau = 0.01, 0.02, 0.1, 0.2, 0.4$; dashed lines from top down: $\tau = 0.6, 0.9$).

To answer this question, recall the ‘fixed-point’ property of the scaling spectrum: if one focuses on a given bin x with $x \ll 1$, then the amount of energy which enters this bin per unit time due to splittings at larger $x' > x$ is exactly equal to the amount of energy which leaves that bin via decays towards smaller $x' < x$. This shows that, via successive branchings, the energy *flows throughout the entire spectrum without accumulating at any value of x* . It therefore accumulates into a ‘condensate’ ($D_{\text{cond}}(x) \propto \delta(x)$) at $x = 0$, according to

$$\mathcal{E}_{\text{flow}}(\tau) = 1 - \int_0^1 dx D(x, \tau) = 1 - e^{-\pi\tau^2}. \quad (6)$$

So far, we have assumed that the evolution remains unchanged down to $x = 0$, but physically this is not the case: when the gluon energies become as low as the typical energy scale in the medium — say, the temperature $T \sim 1$ GeV for a QGP —, then the gluons ‘thermalize’ and disappear from the jet. Thus, our above conclusion about the ‘condensate’ should be more properly formulated as follows: *via successive branchings, the energy flows towards small- x at a rate which is independent of x , and eventually reaches the borderline at $x_{th} \equiv T/E$ between the ‘jet’ and the ‘medium’ at a rate which is independent of the detailed mechanism for thermalization and of the precise value of the medium scale T* . This energy will be recovered in the medium at large angles $\theta \gtrsim \theta_{th}$ w.r.t. the jet axis, with θ_{th} obtained by replacing $x_0 \sim x_{th}$ within the expression for θ_0 shown in Eq. (4).

An energy flow at a rate which is independent of the energy (i.e. uniform in x) is the distinguished signature of *wave turbulence* [14]. This phenomenon is well known in the context of scalar field theories, but it was not expected in the context of QCD, for the following reason: as above discussed, the existence of a turbulent flow requires fine cancellations between ‘gain’ and ‘loss’ terms, which in turn requires the branching process to be *quasi-local in x* . Or, the QCD radiation (bremsstrahlung) is reputed for being highly non-local: all the familiar parton cascades, like DGLAP or BFKL, are dominated by very asymmetric splittings, in which the splitting fraction z is close to zero, or to one. The medium-induced cascade is new in that respect: the splitting of a *soft* gluon, i.e. the process $x \rightarrow zx, (1-z)x$ with $x \ll 1$, is controlled by *quasi-democratic branchings* with $z \sim 1/2$: the offspring gluons carry commensurable fractions of the energy of their parent gluon [8]. This is ultimately related to the peculiar energy dependence of the BDMPSZ spectrum (2), which in turn reflects the LPM effect.

The above picture has interesting consequences for the phenomenology. To see this, let us repeat the exercise in Eq. (4), that is, compute the energy radiated at large angles $\theta > \theta_0$ or, equivalently, small $x < x_0$ (with $x_0 \gg x_{th}$ though), but including the effects of multiple

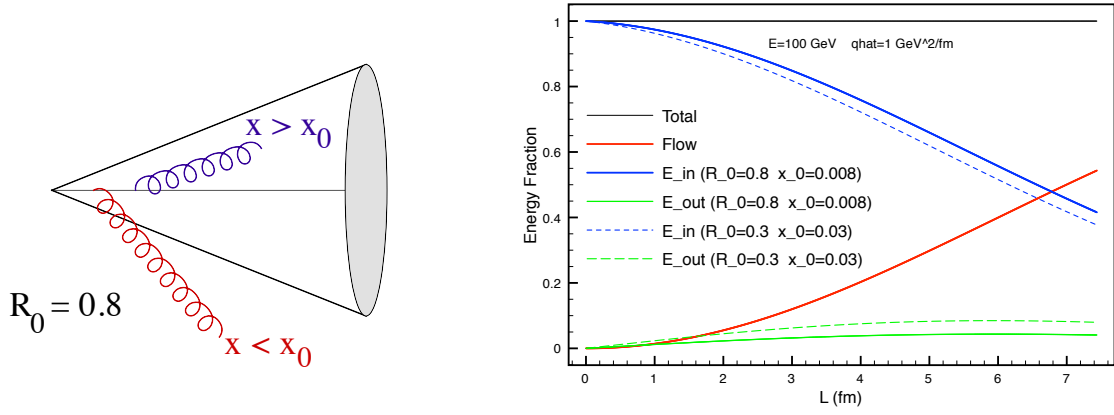


Figure 4: The energy balance between the inner part of a conventionally defined ‘jet’ with angular opening R_0 and the region outside the ‘jet’, plotted as a function of the medium size L , for two values $R_0 = 0.3$ and $R_0 = 0.8$.

branchings. This includes two contributions: the integral over the low- x part of the spectrum (5) at $x \leq x_0$ and the energy (6) carried by the turbulent flow. One finds

$$\mathcal{E}(\theta > \theta_0, \tau) \simeq 2\tau\sqrt{x_0}e^{-\pi\tau^2} + (1 - e^{-\pi\tau^2}) \simeq 2\tau\sqrt{x_0} + \pi\tau^2, \quad (7)$$

where the second approximation holds for $\pi\tau^2 \ll 1$. (Notice that for a jet with $E = 100 \text{ GeV} \approx 2\omega_c$, Eq. (3) implies $\tau \simeq \bar{\alpha} \approx 0.3$.) The flow piece in Eq. (7), which is independent of x_0 , dominates over the non-flow piece for any $x_0 < \tau^2$: *the energy lost by the jet at large angles $\theta > \theta_0$ is predominantly carried by the turbulent flow and hence is independent of θ_0 .*

This is illustrated in Fig. 4 which shows the energy fraction E_{in} inside a jet with angular opening R_0 (E_{in} is the complement of Eq. (7), i.e. the energy contained in the large- x part of the spectrum at $x > x_0$), together with the two components, “Flow” and “Non-flow” (denoted as “ E_{out} ”), of the energy fraction outside the jet, Eq. (7), as functions of the medium size L . As visible there, when increasing the jet angle from $R_0 = 0.3$ and $R_0 = 0.8$, i.e. by almost a factor of 3, the energy captured inside the jet increases only slightly. Moreover, E_{in} represents less than 80% of the total energy for $L \geq 4 \text{ fm}$. The difference (more than 20%) is essentially associated with the flow component, which is independent of R_0 . This picture is in a remarkable agreement with the detailed analysis of the di-jet asymmetry by the CMS collaboration [3].

References

1. **Atlas** Collaboration, G. Aad *et al.*, Phys. Rev. Lett. **105** (2010) 252303.
2. **CMS** Collaboration, S. Chatrchyan *et al.*, Phys. Rev. **C84** (2011) 024906.
3. **CMS** Collaboration, S. Chatrchyan *et al.*, Phys.Lett. **B712** (2012) 176.
4. R. Baier, Y. L. Dokshitzer, A. H. Mueller, S. Peigne, and D. Schiff, Nucl. Phys. **B483** (1997) 291; Nucl. Phys. **B484** (1997) 265.
5. B. G. Zakharov, JETP Lett. **63** (1996) 952.
6. J. Casalderrey-Solana and E. Iancu, JHEP **1108** (2011) 015.
7. J.-P. Blaizot, F. Dominguez, E. Iancu, and Y. Mehtar-Tani, JHEP **01** (2013) 143.
8. J. -P. Blaizot, E. Iancu, and Y. Mehtar-Tani, arXiv:1301.6102 [hep-ph].
9. Y. Mehtar-Tani, C. A. Salgado and K. Tywoniuk, Phys. Lett. B **707** (2012) 156
10. Y. Mehtar-Tani, C. A. Salgado and K. Tywoniuk, JHEP **1210** (2012) 197
11. R. Baier, A. H. Mueller, D. Schiff, and D. Son, Phys.Lett. **B502** (2001) 51.
12. R. Baier, Y. L. Dokshitzer, A. H. Mueller, and D. Schiff, JHEP **09** (2001) 033.
13. S. Jeon, G. D. Moore, Phys.Rev. **C71** (2005) 034901.
14. V.E. Zakharov, V.S. L’vov, and G. Falkovich, *Kolmogorov spectra of turbulence, Vol. 1* (Springer-Verlag, 1992, 264 p.)